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DOI

[10.1103/PhysRevB.107.094105](https://doi.org/10.1103/PhysRevB.107.094105)

Publication date

2023

Document Version

Final published version

Published in

Physical Review B

Citation (APA)

Davidson, J. H., Das, A., Alfasi, N., Cone, R. L., Thiel, C. W., & Tittel, W. (2023). Quadratic Zeeman spectral diffusion of thulium ion population in an yttrium gallium garnet crystal. *Physical Review B*, 107(9), Article 094105. <https://doi.org/10.1103/PhysRevB.107.094105>

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Quadratic Zeeman spectral diffusion of thulium ion population in an yttrium gallium garnet crystal

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(Received 13 October 2022; revised 24 January 2023; accepted 3 February 2023; published 9 March 2023)

The creation of well-understood structures using spectral hole burning is an important task in the use of technologies based on rare-earth ion-doped crystals. We apply a series of different techniques to model and improve the frequency dependent population change in the atomic level structure of thulium yttrium gallium garnet (Tm:YGG). In particular we demonstrate that, at zero applied magnetic field, numerical solutions to frequency-dependent three-level rate equations show good agreement with spectral hole-burning results. This allows us to predict spectral structures given a specific hole-burning sequence, the underpinning spectroscopic material properties, and the relevant laser parameters. This enables us to largely eliminate power-dependent hole broadening through the use of adiabatic hole-burning pulses. Although this system of rate equations shows good agreement at zero field, the addition of a magnetic field results in unexpected spectral diffusion proportional to the induced Tm ion magnetic-dipole moment and average magnetic-field strength, which, through the quadratic Zeeman effect, dominates the optical spectrum over long timescales. Our results allow optimization of the preparation process for spectral structures in a large variety of rare-earth ion-doped materials for quantum memories and other applications.

DOI: [10.1103/PhysRevB.107.094105](https://doi.org/10.1103/PhysRevB.107.094105)

I. INTRODUCTION

Rare-earth ion-doped crystals (REICs) are interesting materials due to their long-lived excited states and their exceptionally long optical coherence times at cryogenic temperatures [1,2]. In particular, along with the possibility for spectral tailoring of their inhomogeneously broadened $4f^N$ - $4f^N$ transitions, this makes them prime candidates for a number of applications in classical and quantum optics. Examples include laser stabilization, rf spectrum analysis, narrow-band spectral filtering, and quantum information storage and processing [2–7].

Thulium-doped yttrium gallium garnet ($\text{Y}_3\text{Ga}_5\text{O}_{12}$, Tm:YGG) is one such material. Its $^3\text{H}_6 \leftrightarrow ^3\text{H}_4$ transition at 795 nm features an optical coherence time of more than 1 ms [8–10], which is one of the longest among all studied REICs. In combination with the accessibility of this transition—within the range of commercial diode lasers—makes it a natural candidate for applications.

The quality of created features and the resulting consequences for associated applications are dependent on the spectroscopic properties of the dopant ions and their numerous interactions with other atomic components in their local crystalline environment [11,12], the details of the optical pumping process, and the spectral and temporal profile of the applied laser pulses [13,14]. Deep understanding of the relation between spectroscopic properties, optical control fields, and spectral diffusion dynamics has resulted in improvements

of this process in a number of other rare-earth-doped materials, including Tm:YAG, Eu:YSO, and Pr:YSO [15–17]. However, this important connection has thus far not been made for Tm:YGG.

In this paper we track the evolution of population within the electronic levels of Tm^{3+} ions in YGG [see Fig. 2(b) for a simplified level scheme] by semicontinuous monitoring of spectral holes for many sequences of applied spectral hole-burning pulses. The characteristic shapes and sizes of these spectral features are matched to a rate equation model that encompasses the ground ($^3\text{H}_6$), excited ($^3\text{H}_4$), and bottleneck ($^3\text{F}_4$) levels in this material with associated lifetimes and branching ratios. At zero magnetic field we see good agreement between our numerical model and measured results across many different pump sequences of varying duration, power, and spectral shape. With the addition of an external magnetic field the agreement with our numerical model disappears as spectral diffusion from local host spins begins to dominate the shape of all spectral features over long timescales. We characterize the nature of this unexpected behavior and expand our model accordingly by adding a spectral diffusion term to account for a quadratic Zeeman interaction with present noisy magnetic fields [18,19].

The paper is structured as follows: In Sec. II we describe the experimental setup used to collect our measurements. In Sec. III we detail the atomic level structure in Tm:YGG and introduce spectral hole burning, the workhorse of our investigations, to select a known set of atomic population. In Sec. IV we introduce and apply a rate equation model which shows good agreement to the measured spectral hole features. In Sec. IV A we detail the use of adiabatic pulses to shape

*These authors contributed equally to this work.

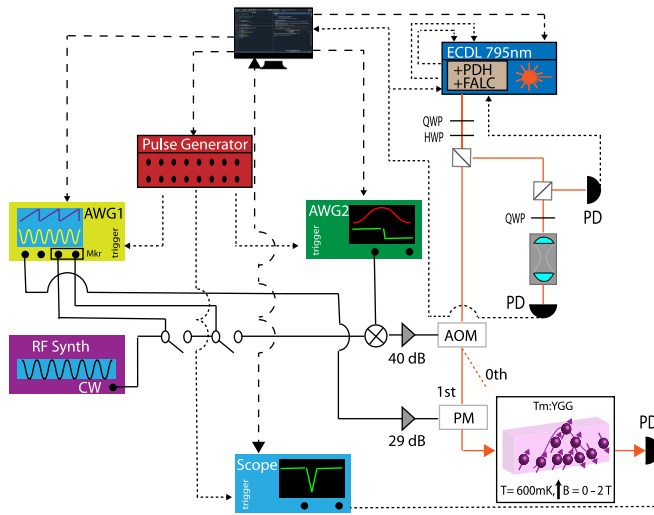


FIG. 1. Schematic of the experimental setup. A PC programs a sequence on a pulse generator (SpinCore Pulseblaster) with nanosecond timing resolution that produces a set of trigger pulses for all devices. When triggered, waveforms written to an arbitrary waveform generator (AWG1, Tektronix AWG 70002A) voltage channel, are subsequently amplified (SHF S126 A) to drive an electro-optic phase modulator (PM) which generates sidebands on the laser light at desired frequencies. The AWG marker channels activate a set of home-built electrical switches which gate the drive signal of an acousto-optic modulator (AOM) to create short pulses from the laser light. This gated AOM signal is mixed with fast arbitrary voltage pulses (AWG2, Tektronix AFG 3102), amplified (Mini-Circuits ZHL-5W-1+), and sent to an acousto-optic modulator (AOM Brimrose 400 MHz) to synthesize controllable-amplitude laser pulses with rise times as short as 100 ns. We use a single light source (Toptica DL Pro 795 nm) for different tasks (optical pumping, pulse generation, etc.). The laser frequency is set via a wave meter (Bristol 871) and locked to a thermally and acoustically isolated high-finesse optical cavity (Stable Laser Systems) via the Pound-Drever-Hall method and a fast feedback loop acting on the laser current and piezo voltage (Toptica PDH and FALC Modules). Transmitted signals from the crystal are directed to a variable-bandwidth photodetector (NewFocus 2051) and displayed on an oscilloscope (Lecroy Waverunner 8100A) configured by the PC and synchronized with the experimental sequence.

spectral holes at zero magnetic field with the goal of creating high-resolution features. Section V shows uncontrolled changes to created spectral holes in the presence of magnetic fields and connects these noise effects to the quadratic Zeeman effect. Section VI extends this quadratic Zeeman connection to the characterization of spectral diffusion results that differentiates the measured results from those predicted by our model over longer timescales.

II. EXPERIMENTAL SETUP

To measure spectral holes, from population storage in various atomic levels, over different timescales in Tm:YGG we use the setup detailed in Fig. 1. A cw diode laser at 795.325 nm is locked to a reference cavity resulting in a linewidth of roughly 5 kHz [20] in the center of the 55 GHz inhomogeneously broadened optical transition of the thulium ions in this material [9]. To craft short pulses of high

extinction ratio, the laser's continuous-wave emission is directed to a free space acousto-optic modulator (AOM). The sinusoidal driving signal of the AOM is mixed with a signal modulated by an arbitrary function generator, which allows programmable control of the transmitted pulse amplitude for the first-order light. After the amplitude control, the pulsed light is directed to a fiber-coupled phase modulator driven using arbitrary waveforms for serrodyne frequency shifting and more complex chirped pulse shapes, as detailed in Sec. IV A. The resulting optical input signals for all experiments in this work are a series of light pulses each with adjustable duration, delay, amplitude, frequency, and frequency chirp. Experimental control of these parameters is handled on a number of different timescales via custom Python scripts that ensures the proper pulses are created and applied at the correct instant [21].

The optical signals are then sent through a polarization controller to a 1% Tm:YGG crystal where all transmitted light is detected by a fiber-coupled photodetector and recorded in real time for subsequent analysis. The crystal sample was grown by Scientific Materials Corp. and is housed in a pulse tube cooled cryostat at 500–700 mK during measurements. A superconducting solenoid centered on the crystal applies a homogeneous magnetic field from 0 to 2 T (using about 1 mA/mT of current) along the crystal's $\langle 111 \rangle$ axis.

III. TM:YGG SITE AND LEVEL STRUCTURE

This series of optical pulses is used to address specific subsets of ions within the garnet crystal host based on the crystal orientation and symmetry, relative to the direction of propagation for the beam, and the external magnetic field above. Garnet crystals such as YGG have cubic crystal structure with O_h^{10} space group symmetry, which yields six Tm^{3+} ion substitution sites, each with D_2 point-group symmetry [9,22–24]. The magnetic and optical behavior of Tm^{3+} ions in each of these sites is identical but the crystal structure leads to effective in-equivalence between the sites due to six different orientations that the ion and its entire local environment can take within the lattice [25]. In essence, the magnetic projection of the field on the ion site determines a number of important properties, including the splittings shown in Fig. 2(b), each ions exact optical transition frequency, and the coherence time of a particular ion. For a few specific directions relative to the crystalline axes pictured in Fig. 2(a), ions at inequivalent sites can be cast into classes that share the same projections of applied electromagnetic fields (\vec{E} , \vec{B}) onto their local site axes.

Given the orientation of our magnetic field, parallel to the crystalline $\vec{B} \parallel \langle 111 \rangle$ axis, we cast the six potential ion sites, shown by the six the parallel pipeds, in Fig. 2(a), into two different classes shown in red and blue. The ions in the two classes experience different magnetic-field projections onto the axes in their local frame. One class features magnetic-field projections along the ion's local X and Z axes (blue), and the other projections along the local Y and Z axes (red). This difference becomes evident from a simple spectral hole-burning experiment, which we use to introduce the remaining results of the paper.

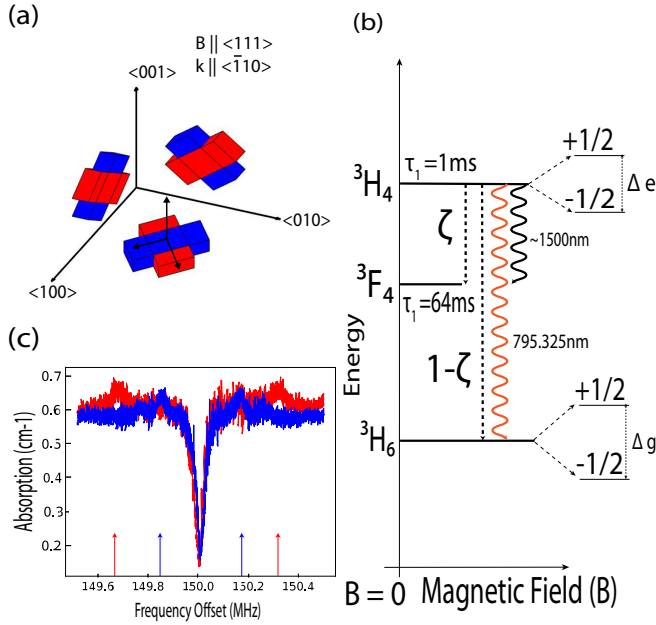


FIG. 2. (a) Depiction of the orientation and symmetry of the six ion substitution sites, pictured as cuboids, relative to the cubic crystal cell axes. For a magnetic field along the crystalline $\langle 111 \rangle$ axis the sites are cast into two classes featuring different field projections, shown in red and blue. Small black arrows indicate the local site X , Y , Z , axes for the red site in the $\langle 100 \rangle$ $\langle 010 \rangle$ plane. The long axis of each cuboid shows the sites local Y axis along which the optical transition dipole is oriented. (b) Relevant energy levels with term symbols for the Tm ion transition of interest with previously measured lifetimes (τ), branching ratios (ζ), and ground and excited Zeeman splittings (Δg , Δe). (c) A pair of hole-burning spectra with the main spectral hole pictured in the center. Each hole, shown in an associated color to panel (a), is burned with a polarization that selects one of the two classes of ions and produces antiholes with different splittings (indicated by the colored arrows). Additional modulations of the optical depth outside the hole originate from a slight orientation offset during crystal cutting and polishing.

In spectral hole burning, a series of optical pump pulses of a chosen duration, power, and frequency excite atomic population in a narrow spectral window within the inhomogeneously broadened absorption spectrum of the ions, in this case centered at a frequency of $795.325\text{ nm} + 150\text{ MHz}$. The excited ions subsequently decay—either back into the original state, or into another energy level [Fig. 2(b) either 3F_4 or $^3H_6 \pm 1/2$] that often belongs to the electronic ground-state manifold. Illuminating the spectral region with a weak laser beam frequency chirped over an interval centered on the frequency of the original pump pulse reveals frequencies of decreased and increased absorption—so-called spectral holes and antiholes [26]. Spectral holes occur at frequencies of reduced ground-state population, i.e., with frequency offset $\delta = 0$ (for the central hole at 150 MHz), $\delta = \Delta e$ (for the side hole), and antiholes can be observed whenever the ground-state population is increased, which happens at $\delta = \Delta g$ and $\delta = \Delta g \pm \Delta e$. Here, Δg and Δe are ground- and excited-state splittings shown in Fig. 2(b). Consult Refs. [13,27,28] for more detailed descriptions of spectral hole burning.

For many systems, the ground- and excited-state splittings depend on the magnitude and direction of the applied magnetic field—which vary for the pair of Tm ion classes [25]. At 7.5 mT we recorded a pair of hole-burning spectra, shown in Fig. 2(c), for orthogonal pumping polarizations. We found two sets of antiholes, each of which split according to the different field projections experienced by the two classes of Tm^{3+} ions. Thus by selecting a particular projection of the burning and readout pulse polarization, we are able to selectively address ions in either of the two created classes, each with different ground and excited splittings. This ability to select out a single class of ions becomes important in Sec. VI as spectral diffusion depends on the magnetic projection for each specific class of ions.

IV. MODELING RESULTS USING THREE-LEVEL RATE EQUATIONS

To further understand the effects of a given hole-burning process on our REIC ensemble we turn to solutions of the Maxwell-Bloch equations that describe the interaction of light with one or many atomic systems [29]. These differential equations can be quite difficult to solve given the complexity that there does not exist a single fixed Rabi frequency to drive all ions [30]. With limited transient coherent effects due to short excitation pulses, small total pulse area, and material inhomogeneities the rate equation approach has been shown to be an effective model [14,31–34]. Thus for the case of Tm:YGG, due to the long T_2 and low optical depth we reduce the Maxwell-Bloch equations to a set of rate equations that describe the conserved total atomic population and how it flows through the different available levels as a function of time [14].

Note that, in the case of narrow-band excitation of an inhomogeneously broadened transition where only a certain portion of atoms are driven, frequency dependence must be added. Following Refs. [14] and [15,35] we describe the dynamics of our atomic ensemble with Eqs. (1)–(3):

$$\frac{\partial n_g(t)}{\partial t} = R(\delta)(n_e - n_g) + \frac{1 - \zeta}{T_e} n_e + \frac{1}{T_b} n_b, \quad (1)$$

$$\frac{\partial n_e(t)}{\partial t} = R(\delta)(n_g - n_e) - \frac{1}{T_e} n_e, \quad (2)$$

$$\frac{\partial n_b(t)}{\partial t} = \frac{\zeta}{T_e} n_e - \frac{1}{T_b} n_b. \quad (3)$$

This system of coupled differential equations describes the relative change of atomic population, $n_{(g,e,b)}(t, \delta)$ in three ion levels as a function of time, and frequency offset, δ . The branching ratio ζ determines how much population decays from the excited state $|e\rangle$ through the bottleneck state $|b\rangle$ before reaching the ground state $|g\rangle$ with respective level lifetimes T_e and T_b . To drive the system, all frequency dependence is contained in the excitation rate $R(\delta)$ which characterizes the applied pulses of light used for hole burning. For Tm:YGG, the level lifetimes and branching ratios are well known at zero magnetic field [9] leaving only one unknown, the frequency dependent optical pumping rate which has to be matched with measured data.

This pumping rate $R(\delta)$ is a function of both the individual ion properties and the applied sequence of laser excitation, or

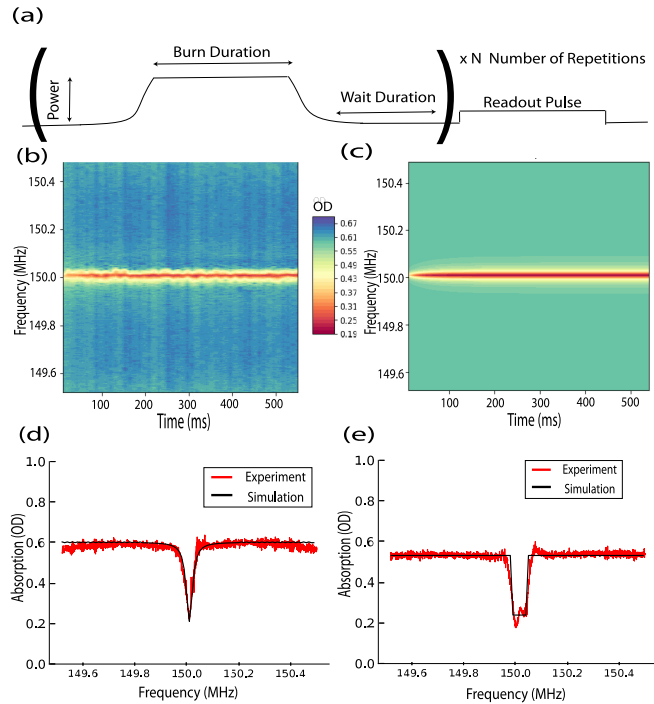


FIG. 3. (a) Simple hole-burning sequence used throughout the measurements. (b) Measured hole evolution. The hole-burning sequence of 1 ms burning pulses, 10 ms waiting periods, and 50 repetitions. The hole deepens as more burn pulses are applied. (c) Evolution of a spectral hole simulated using Eqs. (1)–(3). For both panels (b) and (c) the spectral hole is detuned by 150 MHz from the center of Tm:YGG inhomogeneous broadening and the ground-state population is normalized to the measured optical depth of the material. (d) The final hole of both methods (experiment and simulations). (e) Square-shaped spectral hole generated using hyperbolic secant burn pulses. Modulations at the bottom of the hole are due to imperfections in the burning pulses, and the square line shows comparison to modeled results from a perfectly square $R(\delta)$ function of matching bandwidth.

burning, pulses. Again from Ref. [15],

$$R(\delta) \propto \gamma \otimes |\mathbb{F}(\delta)_{\text{opt}}|^2 \quad (4)$$

is a convolution of the two factors, the homogeneous linewidth γ of the ionic transition and the positive frequency component of the Fourier spectrum of the applied pulse sequence $\mathbb{F}(\delta)_{\text{opt}}$. For the case of Tm:YGG the homogeneous linewidth of the ions is expected to be narrower than 600 Hz [9]. For any $\mathbb{F}(\delta)_{\text{opt}}$ with features greater than the homogeneous linewidth of the ions this convolution is dominated by the Fourier spectrum of the pumping pulse, which is determined by laser and sequence characteristics such as pump power, pulse duration, and pulse spacing, and spectral lineshape.

We perform many different hole-burning experiments with differently timed burning sequences of various powers, durations, and waiting periods in a sequence matching Fig. 3(a) [36]. One example is shown in Fig. 3(b), which shows the measured evolution of a spectral hole after a series of 50 pumping pulses (1 ms) and waiting times (10 ms) where the spectral hole is read out after the every burn pulse to capture the evolution and create each slice of the image.

Then, with these parameters set, and the known lifetimes and branching ratios of Tm:YGG, we numerically solve the rate equations above for all times throughout the same burning sequence after setting the initial conditions before pumping to $n_g = 1$ (i.e., after initializing all population in the ground $^3\text{H}_6$ state). The resulting surface is shown in Fig. 3(c), showing the simulated development of a spectral hole over time for parameters matching Fig. 3(b) and a chosen $R(\delta)$ parameter.

As the optical depth of a measured spectral hole is proportional to the ground-state atomic population left to absorb, we normalize the remaining population using the optical depth before any pumping has occurred and match the results generated by experiment and by numerical simulation for each particular burning sequence [14,15]. Note that since we do not have an analytical function which describes the solution to the system of rate equations, the matching is not a fit that outputs a mean standard error. However, as a “goodness of fit” metric, we calculate the point-wise average difference between the measured and simulated surfaces. Simulation and measurement at the end of this particular full sequence is shown in Figs. 3(d) to demonstrate close agreement for the matched parameters. Discussion of more examples of different burning sequences, including adiabatically shaped pulses [Figs. 3(e)], a discussion of the resulting holes with respect to the level lifetimes, and a note on the power dependence of the laser lineshape are included in the supplementary information.

We consider many different $R(\delta)$ functions to validate our model, only one of which creates the closest match between simulation and results for all the different measured burning sequences. This allows us to uniquely identify the $R(\delta)$ parameter as a property of our laser source and its interaction rate with the ions.

All burning sequences assume a source with a Lorentzian shaped laser line of $\nu = 5$ kHz linewidth and a excitation rate amplitude a , tunable between approximately 300 Hz and 2.5 kHz, via AOM drive power. This yields an $R(\delta)$ function of the form

$$R(\Delta) = a \frac{\nu^2}{\nu^2 + (\delta - \delta_o)^2}$$

$$\text{for } a \in [0.3, 2.5] \text{ kHz,}$$

$$\nu = 0.005 \pm 0.002 \text{ MHz,}$$

$$\delta_o = 150 \pm 0.01 \text{ MHz,}$$

where δ_o is the detuning from the thulium absorption line center.

Thus our model for population change driven by this known optical spectrum allows us to predict the size and shape of any spectral feature resulting from a given burning sequence in this material at zero external magnetic field. This strategy extends to other materials as long as the relevant branching ratios and lifetimes are known.

Adiabatic pulse shaping

To confirm our measurements and exclude spectral hole broadening mechanisms from the optical pumping process, we sought to change the spectral shape of our excitation light and create controllable changes to the width and shape of the resulting spectral holes. We use adiabatic pulse shaping to isolate the shape and width of created spectral features for

further optimization while continuing to match to solutions of our predictive model.

An important method for controlling the shape and width of a desired spectral feature stems from an early goal of NMR to design a pulse which creates highly efficient yet highly selective population inversion [37]. The resulting coherent population inversion over a narrow frequency bandwidth is described by a unique analytical solution to the Maxwell-Bloch equations [38]. This method is known as a hyperbolic secant pulse and relies upon slow adiabatic tuning of the amplitude and phase of the driving pulse. This results in the elimination of unwanted frequency components around the desired bandwidth, and the creation of a pulse with a particularly square spectrum [14,39]. For REICs, hyperbolic secant pulses were first considered for quantum computing in order to create high-fidelity π pulses for narrow bandwidth ensembles of ions [17,40].

The original Maxwell-Bloch solutions call for simultaneous slow modulation of the amplitude and phase, each with a different shape. This is equivalent to modulation of the pulse amplitude and frequency shifting during the duration of the pulse [37]. From a hardware standpoint the problem becomes creating fast electronic control of the amplitude $A(t)$, via AOM voltage, and instantaneous frequency $f(t)$, via phase modulator drive, as in Ref. [16]. To account for the separate elements responsible for each modulation, a variable delay was added to the amplitude modulation to ensure that the correct portion of each pulse is temporally aligned with the corresponding frequency shift.

This method of creating spectral features is in contrast to the common method of linearly swept serrodyne modulation [39,41,42]. For broadband features both methods are fairly similar, but for creating narrow-band spectral features on the order of the pump-laser linewidth, the spectrum of the adiabatic pulse shows much steeper rising and falling edges than that of linearly shifted pulses.

To show the clear difference in spectral hole shape using this method, we carry out the same hole-burning experiment as above with the key difference being that each burning pulse is subject to adiabatic amplitude and frequency modulation. The simulated and measured results of these hole-burning experiments are compared in Figs. 3(f). The hole shape clearly changes corresponding to the altered—squarer—excitation spectrum $R(\delta)$. The shape of the resulting hole spectrum is controllable in terms of width by altering the frequency range and speed of the hyperbolic secant pulse modulations. Although the lower limit of hole width is still related to original laser linewidth, adiabatic modulation suppresses Lorentzian wings of the spectrum, and will allow for added power in the burning process.

V. MAGNETIC NOISE

With the ability to control the frequency chirp range and the amplitude and duration of each burning pulse, creation of many possible spectral features, at zero magnetic field, is well understood from the model in Eqs. (1)–(3). However, the lifetime of the 3F_4 the bottleneck level in Tm:YGG creates an upper limit for burning deep yet sharp holes with a reasonable laser power, see the Supplemental Material [36]. To improve

the depth and lifetime of the resulting features, we now add an applied magnetic field to create a longer-lived bottleneck state. With the addition of a magnetic field the ground and excited states of Tm:YGG each split into spin states with a spin-Hamiltonian given by [25]. The created ground-state spin level has a much longer lifetime than the 3F_4 state previously used as the bottleneck level, prolonging eventual decay to the ground state [10,31].

It is worth noting at this point that $R(\delta)$ is the main term that determines frequency dependence in the model. Thus, with applied magnetic field an identical bandwidth of initial population is excited by the same pulse sequences carried out at zero field. Ideally, only the duration of the subsequent decay path from the excited state can change as magnetic fields are applied leading to deeper longer lived spectral holes. However, it will be seen in the following section(s) that though we can initially create spectral features of a shape matching the zero-field case, quadratic Zeeman coupling between magnetic-field noise and the local spin environment of the Tm ions creates significant changes to these features over long timescales.

A significant portion of electronic population from the 3F_4 level decays to the additional spin level before reaching the original ground state, creating an equivalent three-level system that should be governed by similar rate equations, albeit with extended bottleneck level lifetime [15]. In addition, to ensure that population accumulates in this level, waiting times between burn pulses must be on the order of the 3F_4 state lifetime to prevent population storage in the 3F_4 state from where it can decay to the original ground state much more quickly. Measuring the created spectral holes over these longer burning and waiting sequences with an applied magnetic field, we find important differences compared with the modeling. To assess these differences, spectra in this section and in Sec. VI are taken post-burn sequence on a longer timescale than those presented in the previous sections. More specifically, we measure a spectral hole, created with an initial width of 40 kHz, periodically every 11 seconds for minutes. Each hole is fit to a Lorentzian to determine the center, depth, and width of the spectral feature.

As the magnetic-field strength is increased we observe a clear impact of magnetic-field noise on the width and center frequency of the spectral holes. This is evident from Fig. 4(a). Here the hole is initially centered at the programed frequency and possesses the expected width, but the shape and center proceed to change in an uncontrolled manner due to fluctuations of the applied magnetic field.

To confirm the origin of the measured effect we create a 40 kHz spectral hole at a fixed magnetic field. A short time after burning has completed, we alter the field by a fixed amount of 1.6 mT, the smallest increment possible with our experimental setup, and measure the subsequent changes to the hole. We expect a splitting and shifting [43] from the linear and quadratic Zeeman terms of the ion spin-Hamiltonian as ions change optical transition frequency in the altered field. The results of these measurements are depicted in Figs. 4(b)–4(f). We find that the splitting and shifting of the spectral hole increases notably as the value of the initial magnetic field was increased from 0.1 to 1.5 T for constant field changes of 1.6 mT.

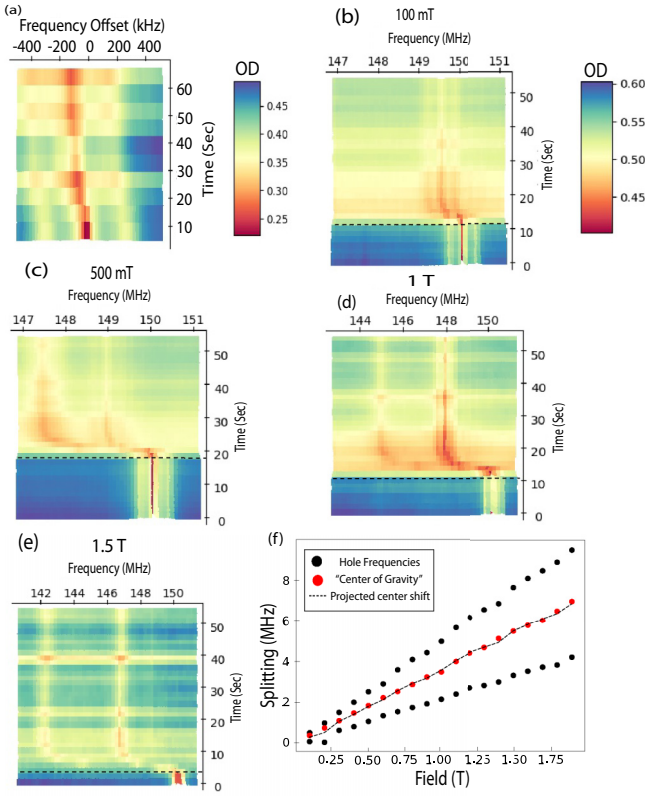


FIG. 4. (a) Spectral hole crafted at 1.8 T, and read out repetitively over the course of one minute. (b)–(e) Shifting and splitting of the hole features for a fixed change of 1.6 mT at initial magnetic fields from 0.1, 0.5, 1.0, 1.5 T. The dashed lines show the moment when the 1.6 mT change in field is made. (f) Hole center frequency shift and splitting into separate hole features after a fixed change of 1.6 mT at initial magnetic fields from 0.1–2.0 T. The dashed line is the center shift projected by the hyperfine tensors for this magnitude of field change. The magnitude of the center shift and the splitting both grow as expected from Ref. [25] for higher absolute fields.

The measured values of the center shift, and splitting of the pair of resulting holes are in good agreement with expectations from the quadratic and linear components of the hyperfine tensor for these ions [25]. The Supplemental Material [36] contains details for calculating the expected shifting from the hyperfine tensors and the applied magnetic-field change. From these measurements, we estimate the field fluctuations in our system to be a factor of 10 smaller than our applied 1.6 mT field change, leading to the shifting of the hole center. The implication is that, at high magnetic fields, noise on the order of μT causes many kilohertz of change in the center hole frequency. This result will prove critical for understanding and creating narrow spectral features at high magnetic fields.

VI. SPECTRAL DIFFUSION

In addition to the central frequency shift, all measured holes are broadened compared with expectations from the burning parameters modeled and tested at zero field, a telltale sign of spectral diffusion [12].

Magnetic-field-dependent spectral diffusion of spins is well understood in the field of rare earths, although it has

more often been considered for the case of Kramers, rather than non-Kramers ions such as Tm [19]. The problem is often cast in terms of interactions in which nearby spins swapping energy with excited members of the ensemble lead to a general loss of coherence and a spreading, or broadening, of spectral features in time [44,45]. Both magnetic field and sample temperature play important roles as they both mediate how much energy is being interchanged directly through spin-spin interaction or through coupling to host phonons [46]. On short-to-medium timescales the characteristic rate and magnitude of these processes are often measured via two-pulse and three-pulse photon echo sequences [19]. For longer timescales spectral hole burning can accomplish a similar task by mapping the change of the hole shape and size as a function of time, magnetic field, and temperature [12].

By adjusting the polarization of the pump light, we first isolate ions with a particular spin splitting at a certain magnetic field, as discussed in Sec. III. Again, we perform spectral hole-burning measurements where burning pulses and waiting periods are timed to guarantee population arrives in the desired spin level. Once hole burning is complete, the hole is read out after varying waiting times using, as above, a weak probe in order to leave the hole unperturbed. We determine (i.e., fit) the widths of all holes for a series of different magnetic-field amplitudes. Width change as a function of time can be described by [19]

$$\Gamma_{\text{hole}} \propto \Gamma_o + \frac{1}{2} \Gamma_{\text{SD}} (1 - e^{-R_s t_{\text{delay}}}). \quad (5)$$

Here, Γ_o is some initial hole width given by $R(\delta)$ of the burning process, which broadens in time. The spectral diffusion approaches a maximum value of Γ_{SD} at a characteristic rate given by R_s . The functional form of the magnetic-field and temperature dependence for Γ_{SD} and R_s can be used to link spectral diffusion to a specific broadening mechanism [19,47,48].

The first portion of evidence for determining the source of the measured diffusion is the different behavior of each magnetic class of ions. Shown in Fig. 5(a) is the hole width of each such class as a function of time for an applied field of 500 mT. For the same external field each class broadens at a different rate. However, when different fields are applied, such that the Zeeman splittings of each class are roughly equivalent, the broadening occurs at the same rate, see Fig. 5(b). This contrast likely rules out contributions of two level systems [31] on this timescale, which are likely to be located randomly throughout the material without an orientation-dependent response. Instead, the diffusion appears to depend on the magnitude of the created magnetic-dipole moment of the Tm ions.

To further isolate the cause of the diffusion we analyze the behavior of a single class of ions as a function of applied field. Utilizing the model in Eq. (5) we fit Γ_{SD} and R_s exclusively for each magnetic class of ions where the functional form of Γ_{SD} is given by [19]

$$\Gamma_{\text{SD}}(B, T) = \Gamma_{\text{max}} \text{sech}^2 \left(\frac{g_{\text{env}} \mu_B B}{kT} \right). \quad (6)$$

Here g_{env} is the g factor of spins in the local crystalline environment, μ_B is the Bohr magneton, B is the applied field, k is the Boltzmann constant, T is the sample temperature, and $\Gamma_{\text{max}}(B)$ is the full width at half maximum of the broadening of the optical transition due to external sources.

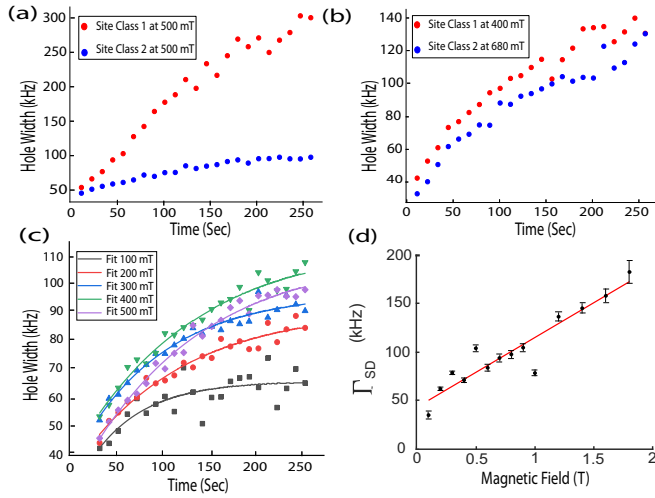


FIG. 5. (a) Spectral hole widths for ions in both magnetic classes as a function of time at 500 mT. The class of ions with the larger magnetic-field projection shows significantly more spectral diffusion. (b) Spectral hole widths for ions in both magnetic classes as a function of time where the applied magnetic field for each class of ions creates roughly equivalent splitting. The spectral diffusion is comparable. (c) Fitted hole width versus time for a single class of ions at a series of different magnetic fields. (d) Linear fit to the variation of Γ_{SD} as a function of magnetic field.

For non-Kramer's ions in general, and all atomic species in the YGG host, since the g factors, g_{env} , are quite small, the thermal distribution of population in various spin levels remains nearly constant for the temperatures and applied fields used in this work, making the sech^2 factor in Eq. (6) effectively constant as well. Any field dependence of Γ_{SD} can then be attributed to the field dependence of $\Gamma_{max}(B)$. As shown in Fig. 5(d) we find Γ_{SD} to increase as the magnetic field increases while R_s remains fairly constant for the measured fields. Inspired by Refs. [19,48] we fit this behavior using a functional form of $\Gamma_{max}(B)$ that relates the optical transition broadening to the quadratic Zeeman effect [36]:

$$\Gamma_{max}(B) = g_j^2 \mu_B^2 B |\Lambda_e - \Lambda_g| B_{Noise} + \Gamma_{max}(0). \quad (7)$$

Here, g_j is the electronic g factor for each level, μ_B the Bohr magneton, Λ_i , $i \in \{g, e\}$ is the hyperfine tensor for the ground and excited states in the transition, B is the applied magnetic field, and B_{Noise} the width of the field distribution around each experimentally set value due to flips of local host spins or external noise. All factors in this expression are known save this empirical B_{Noise} , which is fit to a value of $87 \pm 20 \mu\text{T}$ using the data of Fig. 5(d).

TABLE I. Magnetic fields experienced at an average Tm ion site due to spin flips of the elemental species in the crystal. Estimated distances are in agreement with measured data from YAG [49]. Only the gallium spins match the order of magnitude of the measured magnetic variations.

Species	Concentration	Effective g factor	Average distance (r)	Average field (B)
$^{169}\text{Tm}^{3+}$	1%	0.0077	17 Å	500 nT
^{71}Ga	40%	0.000 71	2.6 Å	15 μT
^{69}Ga	60%	0.000 92	2.6 Å	19 μT
^{89}Y	99%	0.000 14	4 Å	850 nT

Following calculations from Ref. [50] and using a model of the YGG crystal cell from Ref. [51] we estimate the magnetic-field variation at an average Tm ion site given the concentrations, unit-cell dimensions, and corresponding g factors of each host crystal constituent. Table I contains estimations of magnetic fields at a potential Tm ion site due to spin flips from each of the crystal constituents. Considering possible contributors, the low Tm ion concentration of this crystal means the average distance to another dopant thulium is likely too small for Tm-Tm interactions to cause noise of this magnitude. Yttrium, although fully concentrated, is also too weakly magnetic. The most likely explanation is the presence of a fully concentrated bath of local gallium spins, each with a moderate nuclear magnetic moment. An additional contribution from a noisy current supply used to power the superconducting solenoid that creates the magnetic field across the crystal, among other factors stemming from geometrical concerns and heterogeneous spin species [52], may explain the mismatch between our rough estimations of spin-flip noise, and the fitted field variations.

VII. CONCLUSION

We conducted a series of spectroscopic measurements that led to a detailed understanding of how to create high-resolution spectral features using Tm:YGG. The possibility for controlling pump field amplitude, frequency chirp, pulse duration, and pulse power give a number of experimental handles with which to optimize the shape of the desired spectral feature. For zero magnetic field, experimental results were shown to be in good agreement with predictions from a three-level, frequency-dependent, rate-equation model for various series of pumping pulses. However, when applying a similar model to a three-level system that includes a ground-state spin level split by magnetic field, magnetic noise, and spectral diffusion dominate the resulting features over long timescales. We attribute large shifts to each ions optical transition frequency due to quadratic Zeeman effect in combination with time varying magnetic fields as the likely cause of the diffusion. This in depth analysis will aid applications that employ spectral features created using Tm:YGG and other rare-earth-doped crystals such as spectral hole-based laser stabilization, rf frequency analysis, and optical quantum memories.

ACKNOWLEDGMENTS

This material is based in part on research at Montana State University sponsored by Air Force Research Laboratory under Agreement No. FA8750-20-1-1004.

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